

GRAIN GROWTH, OPTICAL POLARISATION AND EXTINCTION IN INTERSTELLAR CLOUDS

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Increases in the wavelength of maximum polarisation and the ratio of total to selective extinction are generally assumed to be the result of grain growth in interstellar clouds (McMillan, 1978; Whittet and Blades, 1980; Clayton and Mathis, 1988). Using a grain model in which the interstellar extinction is explained by amorphous carbon (a-C) and hydrogenated amorphous carbon (a-C:H) mantles on silicate cores (Duley, 1987; Jones, Duley and Williams, 1987; Duley, Jones and Williams, 1988) we have theoretically studied the effects of these absorbing carbonaceous mantles on the polarising properties of large (500 Å - 2500 Å radius) silicate grains. We adopt the polarisation model of Mathis (1986) and show that carbon-coated silicate grains can explain the relationship between the wavelength of maximum polarisation  $(\lambda_{\text{max}})$  and R observed toward dark clouds.

## References

Clayton, G.C. and Mathis, J S.: 1988, Ap. J. 327, 911.

Duley, W.W.: 1987, Mon. Not. R. astr. Soc. 229, 203.

Duley, W.W., Jones, A.P. and Williams, D.A.: 1987, Mon. Not. R. astr. Soc. 229, 213.

Jones, A.P., Duley, W.W. and Williams, D.A.: 1988, Mon. Not. R. astr. Soc. in press.

McMillan, R.S.:1978, Ap. J. 225, 880.

Whittet, D. C. B. and Blades, J. C.: 1980, Mon. Not. R. astr. Soc. 191, 309.